

- 1 a) Prob for one 5 in 1<sup>st</sup> roll:  $(1/6)$   
 b) Prob for 5 followed by anything other than 5 is  $(1/6) \times (5/6)$   
 AND Prob for not-5 followed by 5 is  $(5/6) \times (1/6)$ . So total  
 Prob to get **ONLY** one 5 is sum of these two, or  
 $2 \times (1/6) \times (5/6) = 10/36 = 0.278$   
 c) Prob for at least one 5 will be the same as part (b) with the addition  
 of the Prob for two 5's in two rolls. That latter is  $(1/6)^2 = 1/36$ .  
 The total is then  $10/36 + 1/36 = 11/36 = 0.306$   
 d) In three rolls of the dice there can ONLY be 0 or 1 or 2 or 3 fives.  
 There are no other possibilities. So the addition of each of the  
 possibilities must add to 1.0  
 e) Prob of ONLY one 5 in eight rolls? To get a five on the 1<sup>st</sup> roll and not the  
 next 7 rolls is just Prob of  $(1/6) \times (5/6)^7$ . But there are seven other ways to get  
 only one 5 in eight rolls. The final result will be  
 $8 \times (1/6) \times (5/6)^7 = 0.372$   
 f) There are many ways to have only two 5's turn up in eight rolls. How many?  
 7 ways for the 2<sup>nd</sup> 5 if the 1<sup>st</sup> roll produces a 5  
 + 6 ways for the 3<sup>rd</sup> or 4<sup>th</sup> or ... or 8<sup>th</sup> roll to be 5 if the 1<sup>st</sup> roll is not 5 but the 2<sup>nd</sup>  
 roll is  
 $+5+4+3+2+1 = 28$  which is the same as  $8!/(8-2)! \times (2)!$  From the general formula.  
 Each particular pair will have a Prob of  $(1/6)^2 \times (5/6)^6$ . So the total Prob will be  
 $28 \times (1/6)^2 \times (5/6)^6 = 0.260$

- 2 a) Rate = 0.455 per atom/wk  
 b) ave. out of 35,000 is  $35,000 \times 0.455 = 15930$  decay in one week.  
 c) After 1 week  $(35,000 - 15,930) = 19070$  remain. In next week there will be  
 $19070 \times 0.455 = 8680$  decaying. That leaves  $19070 - 8680 = 10,390$  at the end of the  
 2<sup>nd</sup> week. In the last week  $10,390 \times 0.455 = 4730$  will decay. That leaves roughly  
 $10,390 - 4730 = \underline{5,660}$  left. (That is the same number of significant figures than we  
 started with.)  
 You can use some algebra to do this. If  $N_0$  is the initial number and  $r$  is the rate  
 per atom per year, then in one year there will be  $r \times N_0$  decaying. That means the  
 number remaining in one year is  $N_1 = (1-r) \times N_0$ .  
 By the second year  $r \times N_1$  decayed, which leaves  $N_2 = (1-r) \times N_1 = (1-r)^2 \times N_0$ .  
 Then by three years there are  $N_3 = (1-r)^3 \times N_0$ .  
 With  $r=0.455$  (so  $1-r = 0.545$ ) and  $N_0 = 35,000$  we get  
 $N_3 = (0.545)^3 \times 35,000 = 5666$ .

For problem 3 I first did it with 3 place accuracy and then repeated it with 2 place accuracy. Either way will do for this problem.

$$3a) f = 89.7 \text{ MHz} = 89.7 \times 10^6 \text{ sec}^{-1} = 8.97 \times 10^7 \text{ s}^{-1}$$

$$f\lambda = c \quad \text{So } \lambda = \frac{c}{f} = \frac{3.0 \times 10^8 \text{ m/s}}{8.97 \times 10^7 \text{ s}^{-1}} = 0.334 \text{ m}$$

$$b) 1/1000 \text{ Watts} = 10^{-3} \text{ J/s}$$

$$\begin{aligned} \text{One photon has energy } E &= hf = (6.63 \times 10^{-34} \text{ J}\cdot\text{s}) \times 8.97 \times 10^7 \text{ s}^{-1} \\ &= 5.95 \times 10^{-26} \text{ J} \end{aligned}$$

So the number of photons needed to make up  $10^{-3} \text{ J}$  every second will be  $10^{-3} \text{ J} \div (5.95 \times 10^{-26} \text{ J}) = 1.68 \times 10^{22}$  photons in one second.

$$3.a) f = 90.0 \times 10^6 / \text{sec} = 9.00 \times 10^7 / \text{sec}$$

$$f\lambda = c \text{ (speed of light)} \quad \text{So } \lambda = c / f = \frac{3.00 \times 10^8 \text{ m/sec}}{9.00 \times 10^7 / \text{sec}} = 3.33 \text{ m}$$

$$b) 1/1000 \text{ Watts} = 1/1000 \text{ Joule / sec} = 10^{-3} \text{ J / sec}$$

$$\begin{aligned} \text{but one photon has energy } E &= hf = 6.6 \times 10^{-34} \text{ J sec} \times 9.00 \times 10^7 / \text{sec} \\ &= 6.0 \times 10^{-26} \text{ J} \end{aligned}$$

So number of photons needed to make the  $10^{-3} \text{ J}$  in one second will be  $10^{-3} \text{ J} \div 6.0 \times 10^{-26} \text{ J / photon} = 1.7 \times 10^{22}$  photons

4.a) 510 nm is in the visible region (see the chart in Lecture 6 notes)

b) Wien's Law: peak of the blackbody spectrum,  $\lambda_{peak} \propto 1/T$

So  $\lambda_{peak1} \propto 1/T_1$  and  $\lambda_{peak2} \propto 1/T_2$  with the same proportionality.

$$\text{Then } \lambda_{peak1} / \lambda_{peak2} = (1/T_1) / (1/T_2) = T_2 / T_1$$

$$\text{and } T_2 = T_1 \times (\lambda_{peak1} / \lambda_{peak2}) = 5700 \text{ K} \times (5.10 \times 10^{-7} / 1.06 \times 10^{-3}) = 2.74 \text{ K}$$

5. For  $n = 3, 2, 1$   $E = -1.5 \text{ eV}, -3.4 \text{ eV}, -13.6 \text{ eV}$ .

So  $3 \rightarrow 2$  releases  $(-1.5) - (-3.4) = 1.9 \text{ eV}$  photon

$3 \rightarrow 1$  releases  $(-1.5) - (-13.6) = 12.1 \text{ eV}$  photon

photon  $E = hf$  and  $h = 6.6 \times 10^{-34} \text{ J sec} \div 1.6 \times 10^{-19} \text{ J / eV} = 4.1 \times 10^{-15} \text{ eV sec}$

Then  $f(3 \rightarrow 2) = 1.9 / 4.1 \times 10^{-15} = 4.6 \times 10^{14} / \text{sec}$

$f(3 \rightarrow 1) = 12.1 / 4.1 \times 10^{-15} = 3.0 \times 10^{15} / \text{sec}$

Using  $\lambda = c / f$  gives  $6.5 \times 10^{-7} \text{ m}$  and  $10^{-7} \text{ m}$  for both wavelengths

You can also include  $2 \rightarrow 1$  since the electron that arrives at  $n=1$  could have gone through 2 transitions,  $3 \rightarrow 2$  and  $2 \rightarrow 1$ .

For  $2 \rightarrow 1$  the energy released is  $(-3.4) - (-13.6) = 10.2$  eV photon.

Then  $f(2 \rightarrow 1) = 10.2 / 4.1 \times 10^{-15} = 2.5 \times 10^{15} \text{ sec}^{-1}$

and  $\lambda = 3.0 \times 10^8 / 2.5 \times 10^{15} = 1.2 \times 10^{-7} \text{ m}$